

A Cosmic Microwave Background (CMB) radiation map showing temperature fluctuations across the sky. The colors range from blue (cooler) to red (warmer), with yellow and orange indicating the warmest spots. The map is centered on the Earth and shows a clear dipole anisotropy.

Solving the Universe

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BIGGEST

A deep field image of space, showing a vast field of stars and galaxies. The background is dark, with numerous bright stars of various colors (white, yellow, orange, blue) scattered across the field. Several galaxies are visible, including a prominent spiral galaxy in the upper center and several smaller, more distant galaxies. The overall scene is a rich, multi-colored star field.

Space!

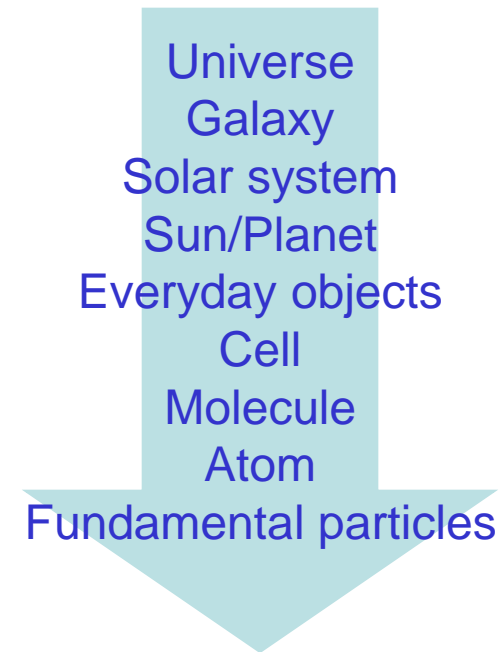
smallest?

Subatomic particles!

Atoms are so small we can't see them --
these particles are even smaller than that!

Opposite Extremes of Size

- Objects in the universe cover a very wide range of sizes
- So-called “fundamental particles” can’t be broken down into smaller pieces
 - Examples include electrons, quarks, and photons, but not protons or neutrons



Why Particle Physics?

- Big Bang: very hot, very dense
- Early universe: only fundamental particles existed
- Higher energy = smaller distances

Looking Back In Time

- Light travels at a fixed speed: 3×10^8 m/s
 - This means the starlight we observe was emitted a long time ago
- Higher energy = looking further back in time
 - We can look all the way back to early stages of the universe – about 380,000 years¹ after the Big Bang!

¹ WMAP Collaboration press release, http://www.nasa.gov/topics/universe/features/wmap_five.html (2008)

Effects of Particle Interactions

- Expansion of the universe
- Formation of atoms
- Formation of galaxies and other structures
- Many other effects, including:
 - Dominance of matter over antimatter
 - Unification of forces
 - Creation/decay of exotic, never-before-seen particles

Cosmic Microwave Background

- CMB radiation
 - Doppler shifted from the visible/near infrared range²
 - Strong evidence for the Big Bang
- Some properties:
 - Very uniform throughout the universe
 - Small variations

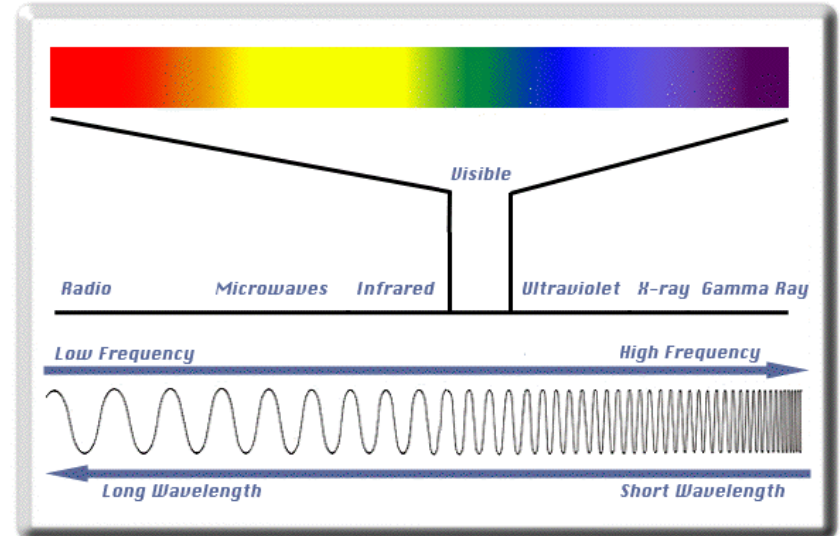
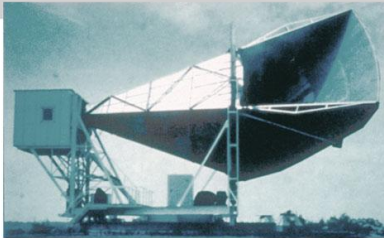


Image courtesy of the University of Minnesota:
http://www.lcse.umn.edu/specs/labs/glossary_items/em_spectrum.html

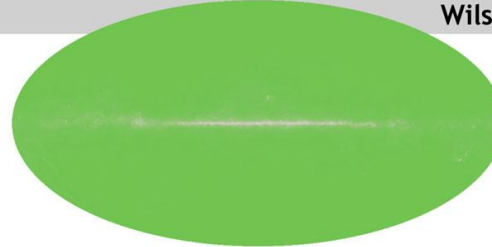
² Ryden, B. *Introduction to Cosmology*. Addison Wesley (2003).

CMB Experiments

1965



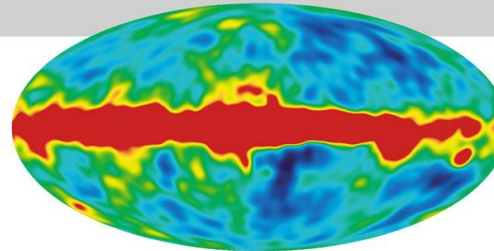
Penzias and
Wilson



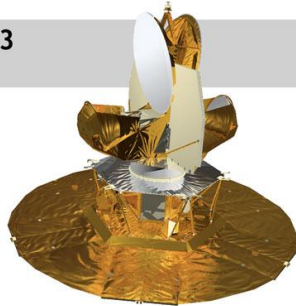
1992



COBE



2003



WMAP

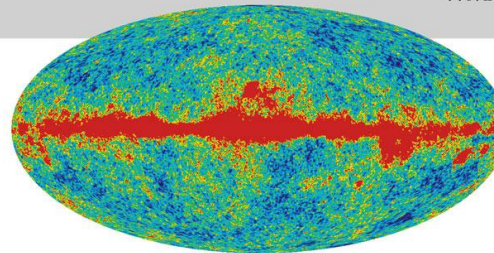


Image courtesy of NASA:
<http://map.gsfc.nasa.gov/media/081031/index.html>

Cosmic Inflation

- Explains why the CMB looks so uniform
- Allows for small variations
- Fixes some of the gaps in the Big Bang theory
- Theories of inflation rely heavily on particle physics

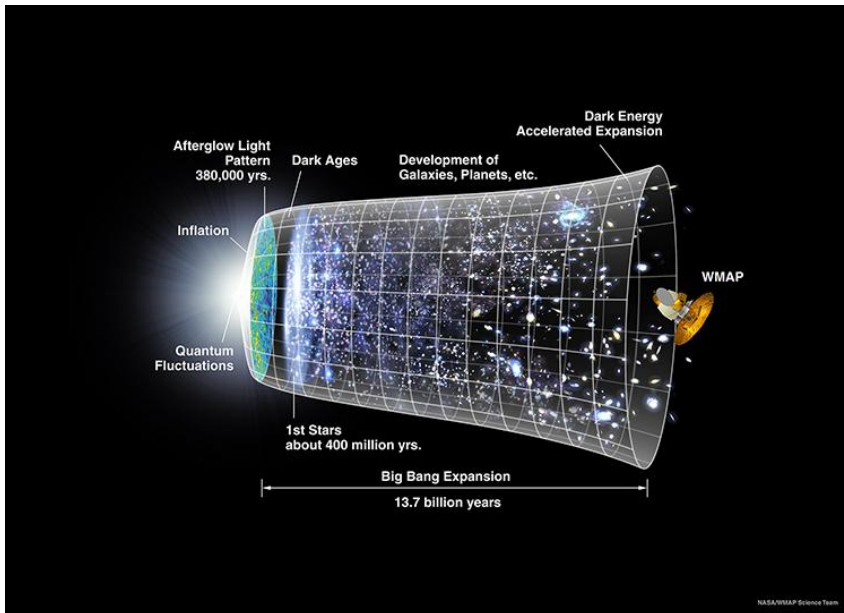


Image courtesy of NASA:
<http://map.gsfc.nasa.gov/media/060915/index.html>

Bringing Inflation into Focus

- So far, experimental evidence allows for a wide range of theories
 - In order to understand how inflation works, we need to narrow this range
- In the end, only experimental measurements can eliminate theories

Just What Is It That I Do?

- Develop and/or modify inflationary models that fit experimental data better
- In order to do this, I:
 - Think of a model that solves the problem in question
 - Perform calculations to obtain estimates of quantities that can be measured by experiments
 - Make predictions of what should happen if the model is true

Questions

